METÓDY RIEŠENIA FYZIKÁLNYCH ÚLOH zima21 – Príklady 5

VZOROVÉ RIEŠENIA

Cvičenie 23.11.2021

Príklad 1

Solution

Apply Gauss' Law to find the E field inside the sphere:

$$\begin{split} \oint \vec{\mathbf{E}} \cdot d\vec{\mathbf{A}} &= \frac{Q_{enc}}{\epsilon_0}, \\ \oint E dA &= \frac{Q}{\epsilon_0} \frac{\frac{4}{3} \pi r^3}{\frac{4}{3} \pi R^3}, \\ E 4 \pi r^2 &= \frac{Q}{\epsilon_0} \frac{r^3}{R^3}, \\ E &= \frac{Q}{4 \pi \epsilon_0} \frac{r}{R^3}. \end{split}$$

Apply circular motion physics,

$$\begin{array}{rcl} m \frac{4\pi^2 r}{T^2} & = & eE, \\ m \frac{4\pi^2 r}{T^2} & = & e \frac{Q}{4\pi\epsilon_0} \frac{r}{R^3}, \\ T^2 & = & \frac{16\pi^3 \epsilon_0 m R^3}{eQ}, \\ T & = & 2\pi \sqrt{\frac{4\pi\epsilon_0 m R^3}{eQ}}. \end{array}$$

Yes, it is independent of r.

Apply Gauss' Law to find the E field outside the sphere:

$$\begin{split} \oint \vec{\mathbf{E}} \cdot d\vec{\mathbf{A}} &= \frac{Q_{enc}}{\epsilon_0}, \\ \oint E dA &= \frac{Q}{\epsilon_0}, \\ E 4\pi r^2 &= \frac{Q}{\epsilon_0}, \end{split}$$

$$E = \frac{Q}{4\pi\epsilon_0} \frac{1}{r^2}.$$

Apply circular motion physics,

$$\begin{array}{rcl} m \frac{4 \pi^2 r}{T^2} & = & eE, \\ m \frac{4 \pi^2 r}{T^2} & = & e \frac{Q}{4 \pi \epsilon_0} \frac{1}{r^2}, \\ T^2 & = & \frac{16 \pi^3 \epsilon_0 m r^3}{eQ}, \\ T & = & 2 \pi \sqrt{\frac{4 \pi \epsilon_0 m r^3}{eQ}}. \end{array}$$

Yes, it is dependent of r. You will hopefully recognize Kepler's law. It is okay to start from a statement like "outside a spherically symmetric charge distribution it is possible to treat the distribution as a point charge."

Use the results of above and find the potential difference between the center and r=2R.

$$\begin{split} \Delta V &= -\int_{2R}^{0} \vec{\mathbf{E}} \cdot d\vec{\mathbf{I}}, \\ &= \int_{2R}^{R} \frac{Q}{4\pi\epsilon_{0}} \frac{1}{r^{2}} + \int_{R}^{0} \frac{Q}{4\pi\epsilon_{0}} \frac{r}{R^{3}}, \\ &= \frac{Q}{4\pi\epsilon_{0}} \left(\frac{-1}{2R} - \frac{-1}{R} + \frac{R^{2}}{2R^{3}} \right), \\ &= \frac{Q}{4\pi\epsilon_{0}R} \end{split}$$

Then use work-energy,

$$\begin{array}{rcl} v & = & \sqrt{\frac{2}{m}e\Delta V}, \\ & = & \sqrt{\frac{2eQ}{4\pi\epsilon_0 mR}}. \end{array}$$

Dimensional analysis is the way to go. a has dimensions of $[L]/[T]^2$, P has dimensions of $[M][L]^2/[T]^3$, c has dimensions of [L]/[T], q has dimensions of [C], and ϵ_0 has dimensions of $[C]^2[T]^2/[M][L]^3$.

Set up an equation such as

$$P = a^{\alpha}c^{\beta}\epsilon_0^{\gamma}q^{\delta}$$

or

$$\left([M][L]^2/[T]^3\right) = \left([L]/[T]^2\right)^\alpha \left([L]/[T]\right)^\beta \left([C]^2[T]^2/[M][L]^3\right)^\gamma \left([C]\right)^\delta$$

Charge is only balanced if $\gamma = -2\delta$. Mass is only balanced if $\gamma = -1$. Similar expressions exist for length and time, yielding

$$P = \frac{1}{6\pi} a^2 c^{-3} \epsilon_0^{-1} q^2$$

The energy radiated away is given by

$$\Delta E = -PT$$
.

where T is determined in the previous sections.

It is possible to compute the actual energy of each orbit, and it is fairly trivial to do for regions r > R, but perhaps there is an easier, more entertaining way. Consider

$$\Delta E = \Delta K + \Delta U$$

and for small changes in r,

$$\frac{\Delta U}{\Delta r} \approx -F = \frac{eQ}{4\pi\epsilon_0} \frac{r}{R^3}.$$

This implies (correctly) that the potential energy increases with increasing r.

$$\frac{\Delta K}{\Delta r} \approx \frac{d}{dr} \left(\frac{1}{2} m v^2 \right) = \frac{1}{2} \frac{d}{dr} \left| r \frac{m v^2}{r} \right|$$

but
$$mv^2/r = F$$
, so

 $\frac{\Delta K}{\Delta r} \approx \frac{1}{2} \frac{d}{dr} \left| rF \right|$

and then

$$\frac{\Delta K}{\Delta r} \approx \frac{eQ}{4\pi\epsilon_0} \frac{r}{R^3}.$$

This implies (correctly) that the kinetic energy increases with increasing r. Not a surprise, since this region acts similar to a multidimensional simple harmonic oscillator. Combine, and

$$\frac{\Delta E}{\Delta r} \approx 2 \frac{eQ}{4\pi\epsilon_0} \frac{r}{R^3} = 2ma$$

Finally,

$$\Delta r = -\left(\frac{1}{6\pi}\frac{a^2}{c^3\epsilon_0}e^2\right)\left(2\pi\sqrt{\frac{4\pi\epsilon_0mR^3}{eQ}}\right)\left(\frac{1}{2ma}\right)$$

This can be simplified, so

$$\begin{split} \Delta r &= -\left(\frac{1}{12\pi} \frac{a}{mc^3 \epsilon_0} e^2\right) \left(2\pi \sqrt{\frac{4\pi \epsilon_0 m R^3}{eQ}}\right), \\ &= -\left(\frac{1}{6} \frac{e^2}{m^2 c^3 \epsilon_0}\right) \left(\sqrt{\frac{4\pi \epsilon_0 m R^3}{eQ}}\right) \left(\frac{eQ}{4\pi \epsilon_0} \frac{r}{R^3}\right), \\ &= -\frac{1}{6} \sqrt{\frac{e^5 Q}{4\pi \epsilon_0^3 R (mc^2)^3}} \frac{r}{R} \end{split}$$

You might want to group these in terms of dimensionless groupings:

$$\Delta r = -\frac{2}{3} \left(\frac{e^2}{4\pi \epsilon_0 Rmc^2} \right) \sqrt{\frac{eQ}{4\pi \epsilon_0 Rmc^2}} r$$

Pick up where we left off, and

$$\frac{\Delta U}{\Delta r} \approx -F = \frac{eQ}{4\pi\epsilon_0} \frac{1}{r^2}$$

This implies (correctly) that the potential energy increases with increasing r.

$$\frac{\Delta K}{\Delta r} \approx \frac{1}{2} \frac{d}{dr} \left| rF \right|$$

$$\frac{\Delta K}{\Delta r} \approx -\frac{eQ}{8\pi\epsilon_0} \frac{1}{r^2}.$$

This implies (correctly) that the kinetic energy decreases with increasing r. Combine, and

$$\frac{\Delta E}{\Delta r} \approx \frac{1}{2} \frac{eQ}{4\pi\epsilon_0} \frac{r}{R^3} = \frac{ma}{2}$$

Follow the same type of substitutions as before, and

$$\Delta r = -\left(\frac{1}{6\pi}\frac{a^2}{c^3\epsilon_0}e^2\right)\left(2\pi\sqrt{\frac{4\pi\epsilon_0 mr^3}{eQ}}\right)\left(\frac{2}{ma}\right)$$

This can be simplified, so

$$\begin{split} \Delta r &= -\left(\frac{1}{3\pi}\frac{a}{mc^{3}\epsilon_{0}}e^{2}\right)\left(2\pi\sqrt{\frac{4\pi\epsilon_{0}mr^{3}}{eQ}}\right), \\ &= -\left(\frac{2}{3}\frac{e^{2}}{m^{2}c^{3}\epsilon_{0}}\right)\left(\sqrt{\frac{4\pi\epsilon_{0}mr^{3}}{eQ}}\right)\left(\frac{eQ}{4\pi\epsilon_{0}}\frac{1}{r^{2}}\right), \\ &= -\frac{1}{3}\sqrt{\frac{e^{5}Q}{4\pi\epsilon_{0}^{3}r(mc^{2})^{3}}} \end{split}$$

You might want to group these in terms of dimensionless groupings:

$$\Delta r = -\frac{4}{3} \left(\frac{e^2}{4\pi\epsilon_0 Rmc^2} \right) \sqrt{\frac{eQ}{4\pi\epsilon_0 Rmc^2}} \frac{R^2}{r}$$

Príklad 2

$$\Phi_{0}(\vec{r}) = \frac{q}{|\vec{r} - a\vec{z}|} + \frac{q}{|\vec{r} + a\vec{z}|} - \frac{2q}{|\vec{r}|}$$

$$= \frac{q}{\sqrt{r^{2} + a^{2} - 2ar\cos\theta}} + \frac{q}{\sqrt{r^{2} + a^{2} + 2ar\cos\theta}} - \frac{2q}{r}$$

$$= \frac{q}{r} \cdot \frac{1}{\sqrt{1 + (a/r)^{2} - 2(a/r)\cos\theta}} + \frac{q}{r} \cdot \frac{1}{\sqrt{1 + (a/r)^{2} + 2(a/r)\cos\theta}} - \frac{2q}{r}$$

$$\approx \frac{q}{r} \left[1 - \frac{1}{2} \left(\frac{a}{r} \right)^{2} + \left(\frac{a}{r} \right) \cos\theta + \frac{3}{8} \left\{ \left(\frac{a}{r} \right)^{2} - 2\frac{a}{r} \cos\theta \right\}^{2} \right]$$

$$+ \frac{q}{r} \left[1 - \frac{1}{2} \left(\frac{a}{r} \right)^{2} - \left(\frac{a}{r} \right) \cos\theta + \frac{3}{8} \left\{ \left(\frac{a}{r} \right)^{2} + 2\frac{a}{r} \cos\theta \right\}^{2} \right] - \frac{2q}{r}$$

$$\approx \frac{q}{r} \left[2 - \left(\frac{a}{r} \right)^{2} + \frac{3}{4} \left\{ 2\frac{a}{r} \cos\theta \right\}^{2} - 2 \right], \text{ to second order in } a/r$$

$$\approx \frac{qa^{2}}{r^{3}} (3\cos^{2}\theta - 1) = \frac{Q}{r^{3}} (3\cos^{2}\theta - 1)$$

(b) On spherical shell $\Phi(r = b) = 0$. Inside shell we have $\Phi_{in}(r) = \Phi_0(r) + \Phi_1(r)$ where $\Phi_1(r)$ satisfies $\nabla^2 \Phi_1 = 0$. Expand potential in Legendre polynomials:

$$\begin{split} &\Phi_{1}(r,\theta) = \sum_{\ell} \left(a_{\ell} r^{\ell} + \frac{b_{\ell}}{r^{\ell+1}} \right) P_{\ell}(\cos \theta) \\ &= a_{0} + \frac{b_{0}}{r} + \left(a_{1} r + \frac{b_{1}}{r^{2}} \right) \cos \theta + \left(a_{2} r^{2} + \frac{b_{2}}{r^{3}} \right) P_{2}(\cos \theta) \end{split}$$

Now at r=0 potential should not diverge, and hence $b_i = 0$. At r=b we have

$$0 = \frac{Q}{b^3} (3\cos^2\theta - 1) + a_0 + a_1 b\cos\theta + a_2 b^2 \frac{(3\cos^2\theta - 1)}{2}, \text{ therefore } a_0 = a_1 = 0$$

and $a_2 = -\frac{2Q}{h^5}$. Therefore the interior potential is given by

$$\Phi_{\rm in}(r) = \Phi_{\rm 0}(r) + \Phi_{\rm 1}(r) = \frac{Q}{r^3} \Big(3\cos^2\theta - 1 \Big) \bigg[1 - \bigg(\frac{r}{b} \bigg)^5 \bigg].$$

Príklad 3

SOLUTION: Solution: The oscillator changes from frequency ω to $\lambda \omega$. In terms of wavefunctions, the initial wavefunction is

$$\psi_i = \left(\frac{m\omega}{\pi\hbar}\right)^{1/4} e^{-\frac{m\omega x^2}{2\hbar}} \tag{1}$$

The new ground state wavefunction is

$$\psi_0 = \left(\frac{m\lambda\omega}{\pi\hbar}\right)^{1/4} e^{-\frac{m\lambda\omega x^2}{2\hbar}} \tag{2}$$

The probability is

$$p = \left| \langle \psi_0 | \psi \rangle \right|^2 = \left| \int_{-\infty}^{\infty} dx \ \psi_0^*(x) \psi_i(x) \right|^2 = \frac{2\sqrt{\lambda}}{1+\lambda}$$
 (3)

It is convenient to reconstruct the ladder operators for the harmonic oscillator. We write

$$a = \frac{x}{\ell} + \frac{i\ell p}{2\hbar} \; ,$$

with arbitrary ℓ . This satisfies $\left[a, a^{\dagger}\right] = 1$, by construction. We furthermore have that

$$a^{\dagger}a + \frac{1}{2} = \frac{x^2}{\ell^2} + \frac{\ell^2 p^2}{4\hbar^2}$$
.

Comparing with H, we equate the ratios of the coefficients of x^2 and p^2 , resulting in $\ell = \sqrt{2\hbar/m\omega}$. The position operator is given by

$$x = \frac{1}{2}\ell(a + a^{\dagger}) .$$

We therefore must calculate

$$\langle \psi_0 | x^4 | \psi_0 \rangle = \frac{\ell^4}{16} \langle 0 | (a + a^{\dagger})^4 | 0 \rangle$$

The operator $(a + a^{\dagger})^4$, when expanded, has $2^4 = 16$ terms. However, only two of these will yield any finite expectation value in the ground state $|0\rangle$. Clearly the only terms which survive must have an a on the left, so as not to annihilate $\langle 0|$, and an a^{\dagger} on the right, so as not to annihilate $|0\rangle$. Furthermore, the number of creation (a^{\dagger}) operators must be the same as the number of annihilation (a) operators. This leaves

$$\begin{split} \left\langle \psi_0 \middle| x^4 \middle| \psi_0 \right\rangle &= \frac{\ell^4}{16} \left\langle 0 \middle| a \, a \, a^\dagger a^\dagger + a \, a^\dagger a \, a^\dagger \middle| 0 \right\rangle \\ &= \frac{3\ell^4}{16} = \frac{3\hbar^2}{4m^2\omega^2} \; . \end{split}$$